

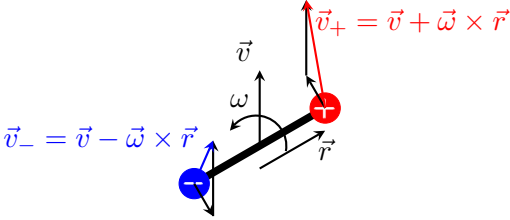
T3: Dipole in a magnetic field

Part (a): Uniform linear motion

Lorentz forces acting on the charges:

$$\begin{aligned}\vec{F}_+ &= q\vec{v}_+ \times \vec{B} = q(\vec{v} + \vec{\omega} \times \vec{r}) \times \vec{B}, \\ \vec{F}_- &= (-q)\vec{v}_- \times \vec{B} = (-q)(\vec{v} - \vec{\omega} \times \vec{r}) \times \vec{B},\end{aligned}$$

where $\vec{r}$ is a vector from the center of mass to the position of the positive charge.



According to Newton's first law, the center-of-mass C of the dipole will move with constant velocity provided that the net force:

$$\vec{F} = \vec{F}_+ + \vec{F}_- = q(\vec{v}_+ - \vec{v}_-) \times \vec{B}, \quad (8)$$

acting on the dipole, is zero. Since $\vec{v}_+, \vec{v}_-$ and $\vec{B}$ are perpendicular, we require $\vec{v}_+ = \vec{v}_-$. It means that dipole does not rotate: $\omega = \omega_0 = 0$.

The pure translation, however, is possible if the pair of forces $\vec{F}_+, \vec{F}_-$, has zero torque about C :

$$\begin{aligned}\vec{\tau} &= \vec{r} \times \vec{F}_+ - \vec{r} \times \vec{F}_- = 2q\vec{r} \times (\vec{v} \times \vec{B}) = \\ &= 2q \left(\vec{v}(\vec{r} \cdot \vec{B}) - \vec{B}(\vec{r} \cdot \vec{v}) \right) = -2q\vec{B}(\vec{r} \cdot \vec{v}).\end{aligned} \quad (9)$$

We conclude that scalar product is zero only when $\vec{v} \perp \vec{r}$, i.e. the initial velocity should be parallel to Y direction.

In summary, the dipole will move uniformly along Y if, and only if, $\vec{v}_0 \parallel Y$ and $\omega_0 = 0$.

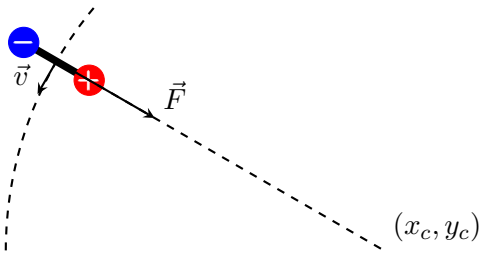
Part (b): Circular motion

The net force can be calculated as:

$$\begin{aligned}\vec{F} &= \vec{F}_+ + \vec{F}_- = 2q(\vec{\omega} \times \vec{r}) \times \vec{B} = \\ &= -2q \left(\vec{\omega}(\vec{B} \cdot \vec{r}) - \vec{r}(\vec{B} \cdot \vec{\omega}) \right) = 2qB\omega\vec{r} = B\omega\vec{p},\end{aligned} \quad (10)$$

where $\vec{p}$ is a dipole moment ($|\vec{p}| = qd = 2qr$ and the direction aligns with $\vec{r}$).

When C orbits a circle, $\vec{F}$ acts as a centripetal force, i.e. it points to the center of the circle. Since $\vec{F} \parallel \vec{p}$, the dipole is always in line with the center of the orbit. Therefore, the orbital angular velocity of C is equal to the angular velocity of rotation of the dipole about C .



The magnitude of the orbital velocity is:

$$v_0 = |\omega_0| R_c$$

From Newton's second law, and accounting that the total mass of the dipole is $2m$:

$$\frac{2mv_0^2}{R_c} = \frac{pBv_0}{R_c},$$

i.e. the magnitude of velocity is:

$$v_0 = \frac{pB}{2m} = \frac{qBd}{2m}$$

and the radius of the orbit is:

$$R_c = \frac{v_0}{|\omega_0|} = \frac{qBd}{2m|\omega_0|}$$

The coordinates of the center of the circle are:

$$(x_c, y_c) = (\pm R_c, 0)$$

where the “+” sign corresponds to $\omega_0 > 0$, i.e. counter-clockwise rotation, and the “−” sign —to clockwise rotation. In either case, the initial velocity should point to the negative Y direction:

$$\vec{v}_0 = -\frac{qdB}{2m}\hat{j}.$$

Part (c): Reversal of the dipole

In (10) we have shown that the net force:

$$\vec{F} = 2q(\vec{\omega} \times \vec{r}) \times \vec{B} = (\vec{\omega} \times \vec{p}) \times \vec{B}.$$

Since the dipole moment $\vec{p}$ rotates with angular velocity $\vec{\omega}$, its time derivative:

$$\frac{d\vec{p}}{dt} = \vec{\omega} \times \vec{p}.$$

From Newton's second law:

$$2m\frac{d\vec{v}}{dt} = \vec{F} = \frac{d\vec{p}}{dt} \times \vec{B}.$$

By integrating the equation, we arrive at an additional conservation law in the system (conservation of the so called “generalized momentum”):

$$2m\vec{v} - \vec{p} \times \vec{B} = \text{const}$$

Thus, if $\vec{p}$ has reversed its direction from $\vec{p}_0$ to $\vec{p}_1 = -\vec{p}_0$, then the velocity:

$$\vec{v}_1 = \vec{v}_0 + \frac{(\vec{p}_1 - \vec{p}_0) \times \vec{B}}{2m} = -\frac{\vec{p}_0 \times \vec{B}}{m}. \quad (11)$$

Since the magnetic field does not perform work on moving electric charges, the kinetic energy of the dipole is conserved:

$$\frac{I}{2}\omega_0^2 = \frac{I}{2}\omega_1^2 + \frac{2m}{2}v_1^2,$$

Here, $I = 2 \times m(d/2)^2 = md^2/2$ is the moment of inertia of the dipole with respect to its center-of-mass. Since v_1 doesn't depend on angular velocities, ω_0 is minimal when $\omega_1 = 0$. Finally,

$$\omega_{\min} = v_1 \sqrt{\frac{2m}{I}} = \frac{p_0 B}{m} \sqrt{\frac{4}{d^2}} = \frac{2qB}{m}$$

Alternatively, we can introduce θ to be the angle between the dipole moment and the axis X ($\theta_0 = 0$) and rewrite the equations of translational motion in coordinates using $\omega = \dot{\theta}$:

$$\dot{v}_x = \dot{\theta} \frac{qBd}{2m} \cos \theta, \quad \dot{v}_y = \dot{\theta} \frac{qBd}{2m} \sin \theta.$$

By integrating these equations, given zero initial velocity, we find how velocity depends on θ :

$$v_x = \frac{qBd}{2m} \sin \theta, \quad v_y = \frac{qBd}{2m} (1 - \cos \theta).$$

Using the expression (9) for the torque, we can write the equation of rotational motion as:

$$I\ddot{\theta} = \tau = -2qB(r_x v_x + r_y v_y) = -\frac{q^2 B^2 d^2}{2m} \sin \theta, \\ \ddot{\theta} + \frac{q^2 B^2}{m^2} \sin \theta = 0, \quad (12)$$

This is the equation of a mathematical pendulum of length L in gravitational field $g = L(qB/m)^2$. And the equivalent question becomes what is the minimal push $\dot{\theta}_0$ required in the bottom position for the pendulum to reach the top position. Kinetic energy of the pendulum $K = \frac{1}{2}mL^2\dot{\theta}_0^2$ will be transferred to the potential energy $U = 2mgL$, from which we find:

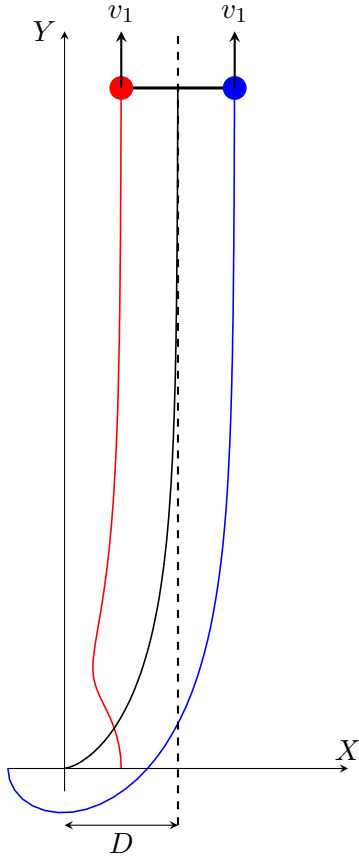
$$\omega_{\min} = \dot{\theta}_0 = \sqrt{4\frac{g}{L}} = 2\frac{qB}{m}.$$

Note. Due to symmetry, both clockwise and counter-clockwise initial rotation with absolute value of $|\omega_0|$ will work.

Part (d): Trajectory asymptote

If dipole's trajectory has an asymptote, then its movement along the asymptote is uniform. Indeed, if there is a linear motion with acceleration, the dipole $\vec{p}$ should be always aligned with the direction of motion, thus, not rotating. and as we found in part (a), the absence of rotation can only be maintained if $\vec{v} = \text{const}$ and $\vec{v} \perp \vec{p}$.

The uniform linear motion requires $\omega = 0$, and this happens in the limit when the orientation is reversed $\vec{p}_1 = -\vec{p}_0$. According to (11), in the limit, the dipole is travelling with the speed $\vec{v}_1 = p_0 B \hat{j}/m$. Thus the asymptote is parallel to Y axis: $x = D$ (for counter-clockwise initial rotation).



If $\vec{R}_+$ and $\vec{R}_-$ are absolute positions of the charges, we can write equation for the angular momentum around the origin L_O :

$$\begin{aligned} \frac{d\vec{L}_O}{dt} &= \vec{R}_+ \times (q\dot{\vec{R}}_+ \times \vec{B}) + \vec{R}_- \times (-q\dot{\vec{R}}_- \times \vec{B}) = \\ &-q\vec{B} \left(\vec{R}_+ \cdot \dot{\vec{R}}_+ - \vec{R}_- \cdot \dot{\vec{R}}_- \right) = -\frac{q\vec{B}}{2} \frac{d}{dt} (R_+^2 - R_-^2). \end{aligned}$$

After integration, we find one more conservation law (conservation of the “generalized angular momentum”):

$$\begin{aligned} \vec{L}_O + \frac{q\vec{B}}{2} (R_+^2 - R_-^2) &= \vec{L}_O + \frac{q\vec{B}}{2} \left((\vec{R}_+ + \vec{R}_-) \cdot (\vec{R}_+ - \vec{R}_-) \right) \\ &= \vec{L}_O + \vec{B}(\vec{R} \cdot \vec{p}) = \text{const}, \end{aligned}$$

where $\vec{R} = \frac{1}{2}(\vec{R}_+ + \vec{R}_-)$ is the position of center of mass. We also used the fact that $q(\vec{R}_+ - \vec{R}_-) = 2q\vec{r} = \vec{p}$.

Initially, centre of mass coincides with origin ($\vec{R}_0 = 0$):

$$L_O(0) = I\omega_0 = 2m \frac{d^2}{4} 2 \frac{qB}{m} = qBd^2. \quad (13)$$

At asymptote, the dipole has reversed direction $\vec{p}_1 = -\vec{p}_0$ and charges are travelling along parallel lines $x = D \pm r$ with the velocity $\vec{v}_1$:

$$\begin{aligned} L_O(\infty) + B(\vec{R}_1 \cdot \vec{p}_1) &= m(D-r)v_1 + m(D+r)v_1 - BDp_0 \\ &= 2mD \frac{p_0 B}{m} - BDp_0 = BDp_0 = BDqd. \quad (14) \end{aligned}$$

Since (13) equals (14), we conclude that $D = d$.

We can arrive to the same conclusion differently. Notice that we are interested in the x coordinate of C at infinity:

$$D = x_{\infty} = \int_0^{\infty} v_x dt = \frac{qBd}{2m} \int_0^{\infty} \sin \theta dt.$$

From (12), we can express $\sin \theta$:

$$\begin{aligned} \int_0^{\infty} \sin \theta dt &= -\frac{m^2}{q^2 B^2} \int_0^{\infty} \ddot{\theta} dt = \\ &= -\frac{m^2}{q^2 B^2} (\dot{\theta}_1 - \dot{\theta}_0) = \frac{m^2}{q^2 B^2} \omega_{\min} = \frac{2m}{qB}. \end{aligned}$$

Finally,

$$D = \frac{qBd}{2m} \frac{2m}{qB} = d.$$

Note. If initial rotation is clockwise ($\omega_0 < 0$), the asymptote has an equation $x = -D$, but the distance to the origin remains the same.

Marking scheme

Part (a): Uniform linear motion		pts
a1	Rationalizes that the net force on the dipole is zero if the two poles move with equal velocities; Just argument $v = \text{const} \Rightarrow \sum \vec{F} = 0$ is 0 pts.	0.7
a2	Concludes that $\omega_0 = 0$.	0.3
a3	Using the argument of zero torque, concludes that the velocity should be perpendicular to the dipole; Just argument $\omega = \text{const} = 0 \Rightarrow \vec{\tau} = 0$: 0.4 pts	0.7
a4	States explicitly that $\vec{v}_0 \parallel Y$ (or $\perp X$).	0.3
Total number of points for part (a)		2.0
Part (b): Circular motion		pts
b1	Derives expression for the magnitude of the net force on the dipole in terms of ω AND states explicitly that it is parallel to the dipole axis OR derives one single vector expression.	0.9
b2	Realizes (drawing or explicit statement) that $\vec{F}$ and the dipole axis point to the center of the orbit, and concludes that ω_0 is equal to the orbital angular velocity.	0.5
b3	Writes down Newton's second law for the circular motion.	0.5
b4	Makes use of the relation $v_0 = \omega R_c$.	0.2
b5	Derives expression for v_0 and specifies its direction (drawing or statement) OR derives one single vector expression for $\vec{v}_0$; if direction is wrong or missing 0.2 pts	0.3
b6	Derives explicitly $R_c = qbD/(2m \omega_0)$. If $ \cdot $ is omitted, still full points.	0.3
b7	Writes down the coordinates of the center of the orbit; 0.2 for correct x_c (including sign), 0.1 for correct y_c ; $x_c = qbD/(2m\omega_0)$ is a correct answer	0.3
Total number of points for part (b)		3.0

Only one of the grading tables should be used for part (c), the one which results in a higher score.

Part (c): Reversal of the dipole		pts
c1	By integrating the equation(s) of motion derives a “generalized momentum” conservation law – a relationship between the linear momentum $2m\vec{v}$ and the dipole moment $\vec{p}$ – in vector form OR for the Cartesian components.	1.5
c2	States explicitly that the kinetic energy of the dipole conserves.	0.3
c3	Writes down explicit expression for the kinetic energy in terms of angular velocity and linear velocity of the center of mass.	0.5
c4	Realizes that ω_0 is minimal when $\omega_1 = 0$ in the reversed position.	0.2
c5	By using the “generalized momentum” conservation, derives explicit expression for the linear velocity v_1 .	0.5
c6	Applies the conservation of energy to find relationship between v_1 and $\omega_{\min}$	0.8
c7	Derives the final expression for $\omega_{\min}$	0.2
Total number of points for part (c)		4.0

Alternative approach: pendulum analogy

Part (c): Reversal of the dipole		pts
c1	Derives the expression $\tau = -B(\vec{p} \cdot \vec{v})$ for the torque. Even if the derivation has been made in parts (a) or (b), the points should be assigned to Task (c); If term $(\vec{B} \cdot \vec{p})$ is not cancelled, still full points	0.5
c2	By integrating the equations of motion, expresses v_x and v_y in terms of θ .	1.5
c3	Writes down the equation of rotational motion in terms of $\sin \theta$.	0.5
c4	States that the angular dynamics of the dipole is equivalent to a large-amplitude oscillation of a mathematical pendulum.	0.3
c5	Realizes that ω_0 is minimal when $\omega_1 = 0$ in the reversed position.	0.2
c6	Applies the conservation of energy to the “equivalent pendulum”.	0.8
c7	Derives the final expression for $\omega_{\min}$	0.2
Total number of points for part (c)		4.0

Part (d): Trajectory asymptote		pts
d1	Rationalizes that the asymptote is parallel to Y , i.e. $x = \pm D$.	0.1
d2	Rationalizes that asymptotically the motion is linear uniform	0.2
d3	Either finds conservation law $\vec{L}_O + \vec{B}(\vec{R} \cdot \vec{p})$ OR writes x_∞ as integral of v_x (with explicit expression for v_x) as a method to find D .	0.3
d4	Correctly computes generalized angular momentum at 0 and ∞ OR uses $\sin \theta \propto \dot{\theta}$ in integral.	0.2
d5	Concludes that $D = d$.	0.2
Total number of points for part (d)		1.0